



Amplitudes and Backgrounds in Split Signature

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YTF 20, 16/12/2020



Motivation

- Classical gravity from amplitudes Δp^μ , Δs^μ , $F_{\text{rad}}^{\mu\nu}$, $\Psi_{\alpha\beta\gamma\delta}$

KMOC [1811.10950]



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BCFW [0501052]

- Classical/quantum double copy

BCJ [0805.3993]

Radiation in EM and gravity



$$F_{\mu\nu} \rightarrow \phi_{\alpha\beta} = F_{\mu\nu} \sigma^{\mu\nu}{}_{\alpha\beta} = (\mathbf{E} - i\mathbf{B}) \cdot \boldsymbol{\sigma}$$

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$$W_{\mu\nu\rho\sigma} \rightarrow \Psi_{\alpha\beta\gamma\delta} = W_{\mu\nu\rho\sigma} \sigma^{\mu\nu}{}_{\alpha\beta} \sigma^{\rho\sigma}{}_{\gamma\delta}$$

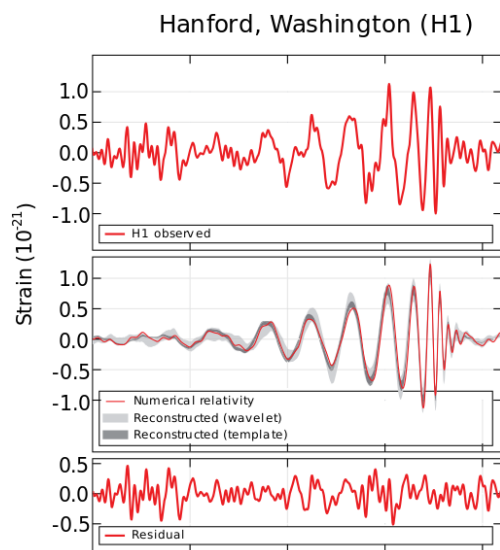
NP basis and amplitudes



$$\Psi = \Psi_0|n\rangle^4 + \Psi_1|k\rangle|n\rangle^3 + \Psi_2|k\rangle^2|n\rangle^2 + \Psi_3|k\rangle^3|n\rangle + \Psi_4|k\rangle^4$$

NP basis and amplitudes

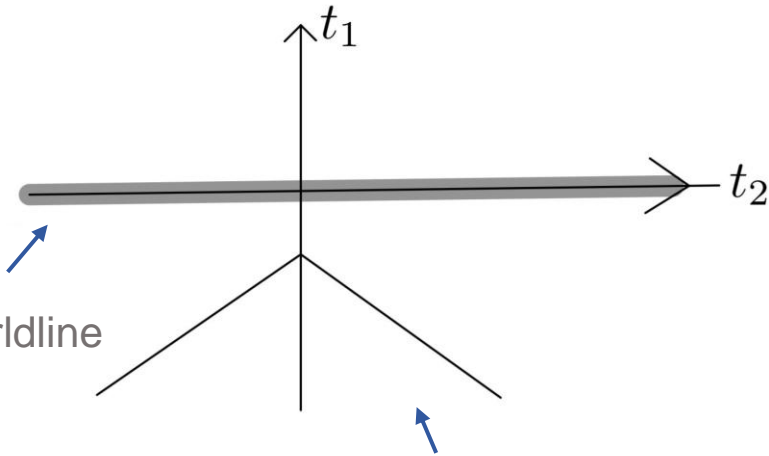
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“knows” about asymptotic evolution,
can compute it with amplitudes!

LIGO collaboration
[1602.03837]

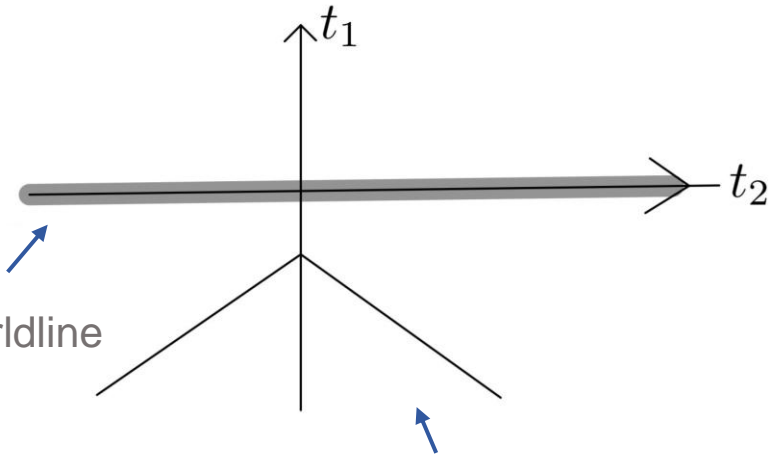
Asymptotic radiation in (2,2)



Massive scalar particle worldline
with $u^\mu = \frac{p^\mu}{m} = (0, 1, 0, 0)$

Vacuum: $\langle \psi | \Psi(x) | \psi \rangle = 0$

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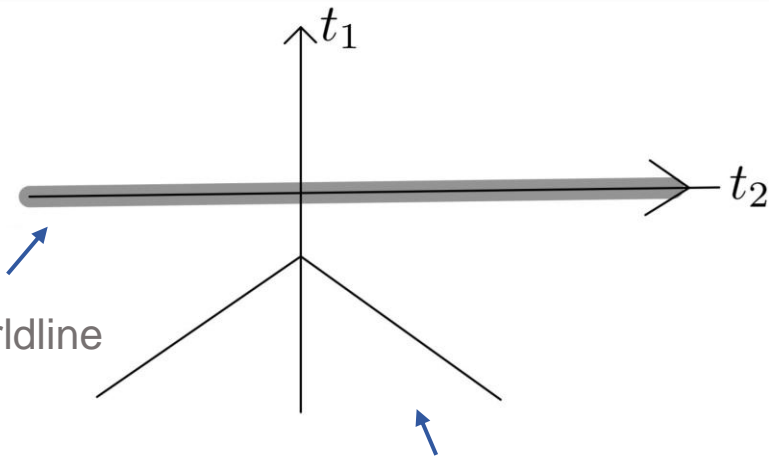


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Measure $\langle \psi | S^\dagger \Psi(x) S | \psi \rangle$ at large t_1 infinity

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$$S = \lim_{t_1 \rightarrow \infty} U(-t_1, t_1)$$

Measure $\langle \psi | S^\dagger \Psi(x) S | \psi \rangle$ at large t_1 infinity

$$|\psi\rangle \sim \int \varphi(p) |p\rangle$$

$\mathcal{M}^{(3)}$ Is enough classically ($\hbar \rightarrow 0$)



- Find the coherent scattered state $\mathcal{M}_h^{(3)} = -\kappa m^2 (u \cdot \varepsilon_h(k))^2$

$$S|\psi\rangle = \exp \left(\sum_h \int d\Phi(k) \delta(2p \cdot k) i\mathcal{M}_{-h}^{(3)}(k) a_h^\dagger(k) \right) |0\rangle$$

A blue arrow points from the $\mathcal{M}_h^{(3)}$ term in the list above to the $\mathcal{M}_{-h}^{(3)}(k)$ term in the exponent of the equation below.

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- Can also compute asymptotic Weyl spinor

$$\langle \Psi(x) \rangle = 2\kappa \text{Re} \int d\Phi(k) \delta(2p \cdot k) i\mathcal{M}_+^{(3)}(k) |k\rangle^4 e^{-ik \cdot x}$$

$$\kappa = \sqrt{32\pi G_N}$$

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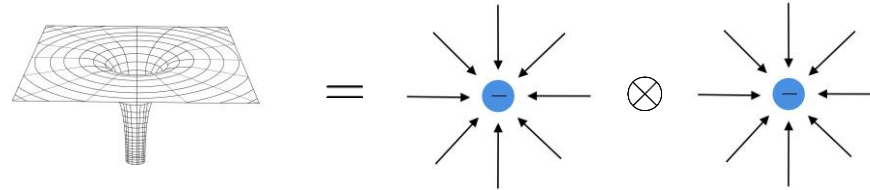
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$$\mathcal{M}^{(3)} \leftrightarrow (\mathcal{A}^{(3)})^2 \sim \langle \phi(x) \rangle$$

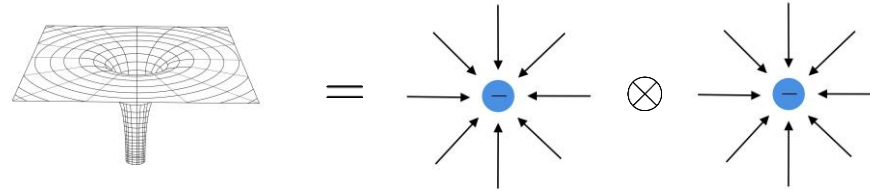
The Classical double copy



$$A^\mu(x) = \phi(x) K^\mu(x), \quad K^2 = 0, \quad K \cdot \partial K^\mu = 0,$$

$$\Rightarrow g_{\mu\nu}(x) = \eta_{\mu\nu} + h_{\mu\nu}(x) = \eta_{\mu\nu} + \phi K_\mu K_\nu,$$

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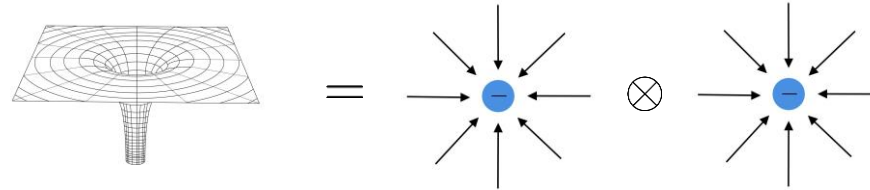
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$$A^\mu(x) \sim \frac{1}{r} K^\mu, \quad K^\mu = \left(1, \frac{\vec{x}}{r}\right) \Rightarrow ds^2 = g_{\mu\nu}^{\text{Sch}} dx^\mu dx^\nu,$$

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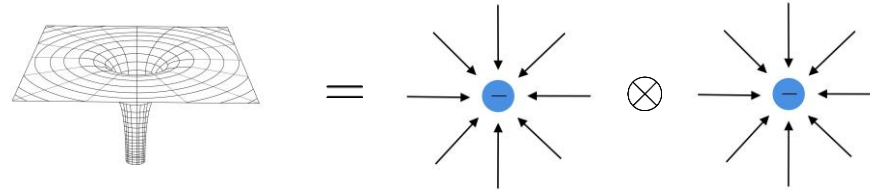
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\Downarrow (2,2)

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Fully characterized
by three point
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Conclusions and future directions



(2,2) can help us understand many things!



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- Simple setting, only 3-point amplitudes
- Close link between background radiation fields and scattering amplitudes
- Classical solutions in GR fully characterised
- Analytically continue back to (3,1)
- To do next: Spin, Taub-NUT..

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Thank you!