# Pushing light millicharged dark matter to its limits [2102.08394] with Joerg Jaeckel

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# Pushing light millicharged dark matter to its limits The very light DM paradigm



 $m \lesssim \mathrm{eV}$ 

# SM

# SM



# Pushing light millicharged dark matter to its limits The very light DM paradigm



 $m \lesssim \mathrm{eV}$ 





# Producing very light dark matter The misalignment mechanism

$$V\left(\varphi\right) = \frac{m^2}{2}\varphi^2$$

Harmonic oscillations

$$\varphi(t) = \varphi_0 \left(\frac{a_0}{a}\right)^{\frac{3}{2}} \cos\left(m\left(t - t_0\right)\right)$$
$$= \Phi(t) \cos\left(m\left(t - t_0\right)\right)$$

DM energy density  $\langle \rho \rangle \sim \frac{1}{2} m^2 \Phi^2(t) \propto a^{-3}$  [Preskill, Wise, Wilczek; 1983]

[Abbott, Sikivie; 1983]

[Nelson, Scholtz; 2011]

[Arias, Cadamuro, Goodsell, Jaeckel, Redondo, Ringwald, 2012]

#### $\rightsquigarrow \quad \ddot{\varphi} + 3H\dot{\varphi} + m^2\varphi = 0$





# Resonant depletion of millicharged DM

**Resonant depletion of millicharged DM** Scalar QED as a toy theory

$$\mathcal{L} = -\frac{1}{4}F^2 + (1)$$

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Annihilation channels are open

$$\phi\phi \to AA$$

Electric charge q should be tiny

$$\ddot{A} + H\dot{A} + \left(\frac{k^2}{a^2} + q^2\varphi^2\right)A = 0$$

 $(D_{\mu}\phi)^{\dagger} D^{\mu}\phi - m^{2}\phi^{\dagger}\phi$ 



#### Parametric resonance Instabilities of the Mathieu equation

Mathieu equation



Solutions have an exponential part



#### **Parametric resonance** Instabilities of the Mathieu equation

Mode functions determine occupation number

$$n_k = \frac{\omega_k}{2} \left( \frac{|\dot{A}|^2}{\omega_k^2} + |A|^2 \right)$$

Unstable solutions = rapid productions of photons  $\forall$  10

$$n_k \propto \exp\left(2\mu_k m t\right)$$





# What about the expansion of the Universe?

# Including the expansion of the Universe

Photons get redshifted

$$\frac{\delta k}{k} \simeq H \delta t$$

Trajectory in the instability chart  $(\mathcal{A}_k(t), \mathcal{Q}(t))$ 

Exponential growth vs redshift

$$\delta t_{\rm exp} \simeq \frac{1}{2\mu_k m}$$





# Including the expansion of the Universe

In order to avoid a fragmentation of the DM

$$\frac{1}{2\mu_k m} \gtrsim \frac{1}{H} \frac{\delta k}{k}$$

Can be translated into a condition on the charge

Consider different regimes

$$\mathcal{Q} \ll 1 \qquad \qquad \mathcal{Q} \gg 1$$





# Including the expansion of the Universe

Narrow resonance

$$\log\left(\frac{1}{n_0}\sqrt{\frac{m}{H}}\frac{\Phi}{qm}\right) \gtrsim \frac{m}{H}\left(\frac{q\Phi}{2m}\right)$$

Broad resonance

$$\log\left(\frac{1}{n_0}\sqrt{\frac{m}{H}}\frac{\kappa}{q^2}\right) \gtrsim \frac{2\kappa}{3}\frac{m}{H}\left(\frac{q\Phi}{2m}\right)$$





# The plasma mass of the photon... ...blocks the resonant depletion

Mathieu parameter shift

$$\mathcal{A}_k \propto \frac{k^2}{a^2} \to \frac{k^2}{a^2} + m_\gamma^2$$

Resonance becomes inefficient

$$m_{\gamma}^2 \gtrsim m^2$$

Plasma mass  $m_\gamma$  [eV]



# **Resonant depletion of millicharged DM** Scalar QED toy theory

Evaluate at the earliest possible time

Massive photons possible

$$q\gtrsim 4\frac{m_{\gamma}^2}{m\Phi}$$



# **Resonant depletion of millicharged DM** Kinetic mixing with a hidden photon



 $q_{\rm eff} = \epsilon q g$ 

# **Resonant depletion of millicharged DM Kinetic mixing case**

DM largely annihilates into HP

Backreaction

$$\rho_A \rightleftharpoons \rho_\phi$$

Coherence

$$\Delta k \gtrsim m v_{mr} \left(\frac{a_{mr}}{a}\right)$$





## **Resonant depletion of millicharged DM** Independent of mixing parameter







## Pushling light millicharged DM to its limits Conclusions

- High occupation number may lead to rapid production of gauge bosons  $n_k \propto \exp\left(2\mu_k m t\right)$
- DM field transfers energy exponentially

Stability puts strong constraints on coupling

Why is there nothing left?



